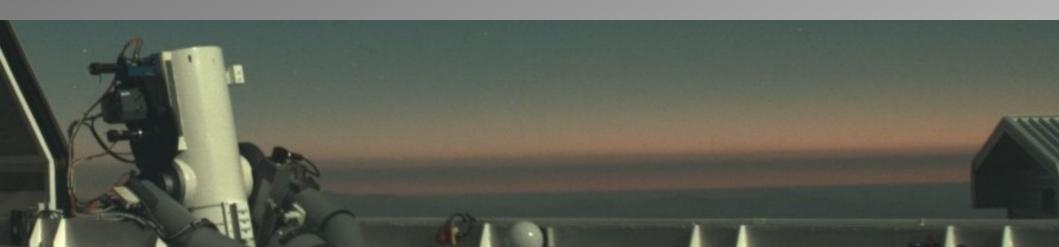
# The GEOS RR Lyr database and survey A professional-amateur project

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## **GEOS**: http://geos.upv.es/

- Created in 1974, already European and pro-am.
- Promote research among amateur astronomers in Europe
- The basic idea is that amateurs should themselves extract scientific information from their observations and publish their results.
- In 1999: creation of a RR Lyr star maxima database.
- During the past years the study of RR Lyr stars has become the main field of interest of GEOS but other variable stars are also studied: eclipsing binaries, δ Scuti stars, Cepheids ... essentially "fast" varying stars

#### References:

Hambsch, F.-J.; Le Borgne, J.-F.; Poretti, E.; GEOS association, 2012, The Journal of the American Association of Variable Star Observers, vol. 40, no. 1, p. 177

Boninsegna, R.; Vandenbroere, J.; Le Borgne, J. F.; GEOS Team, 2002, ASP Conference Proceedings, Vol. 259. 166

## **GEOS RR Lyr Database**

A tool for the study of RR Lyr stars

http://rr-lyr.irap.omp.eu/dbrr/

Allow to follow period variations since star's discovery, some times over 100 years ago.

Help to preparation of observations of RR Lyr stars,

## **GEOS RR Lyr Survey**

Add significantly more maximum timings of the brightest RR Lyr stars essentially using robotic telescopes.

Study fainter understudied stars to refine their period and find new stars which exhibit Blazhko effect.

Characterize the Blazhko effect, and other long term variations of RR Lyr stars.

## **GEOS RR Lyr Database**

## Aim of the database: Maintaining an up-to-date list of observed maxima of RR Lyr stars.

Starting idea in 1999: updating the data published in the book "RR Lyr Stars" written by Vladimir P. Tsesevich in the years 1960's (1966, 1969 for the English translation).

And make it available on the web,

First, collecting published historical times of maximum made with the efficient help of Anton Paschke, Massimiliano Martignoni and Francesco Acerbi in the early years of the project.

We make a survey of recent publications periodically

A couple of days ago (2019 September 10): 105068 maximums on 4060 stars

To access data: a **web interface** to the database list of maxima, O-C curves, basic data of the stars

The web site is hosted by IRAP.

## GEOS RR Lyr Database: 1999-2019

### 3<sup>rd</sup> version of web interface:

- Interface coded in html + php + css
- Web pages created dynamically
- Maximum data in database mysql
- Indexed on unique star identification
- Uses alias names



#### **Database tables**

- Catalog: coordinates, type, magnitudes
- Alias names
- Elements
- Maximums
- Bibliographical references
- Comments

#### Web interface

Access to data:

- Access by constellation and GCVS names
- Access by star name
- Complete list of stars with maxima
- Access by coordinates / Access to catalog
  - Center of field or field limits
  - Center on given star

#### Access to data

- · Complete list of stars with maxima
- Access by constellation and GCVS names
- · Access by coordinates
- · Access by star name



- · List observers (alphabetic) (score)
- Bibliography
- Statistics on RR Lyr maxima
- The Blazhko star catalog
- RR Lyr star spectroscopy

#### Database numbers

#### Date of last updates

General RR Lyr catalog: 2019-06-09 17:01:36

(62313 entries)

Element table: 2019-08-19 18:19:18

(19159 entries)

Maximum table: 2019-09-06 18:44:03

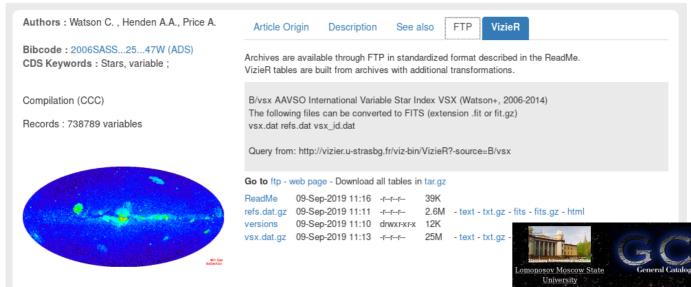
105068 maximums on 4060 stars

## Making of the RR Lyr catalog

Access to GCVS and VSX catalogs

VSX catalog is updated every Monday at CDS:

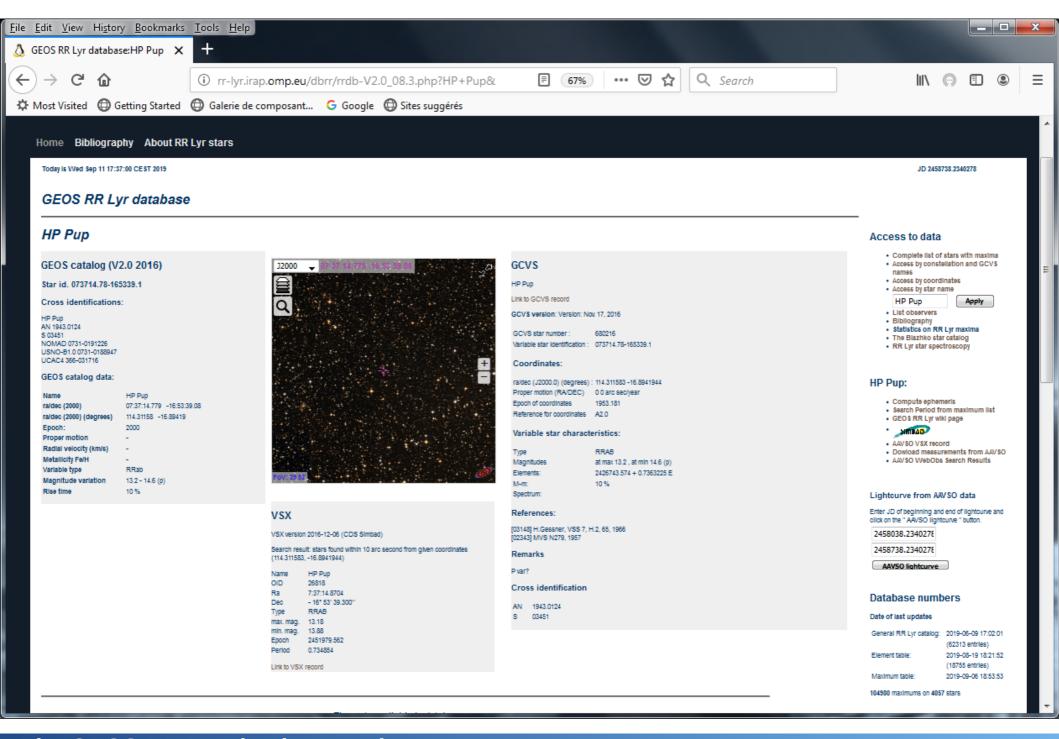
http://cdsarc.u-strasbg.fr/viz-bin/cat/B/vsx

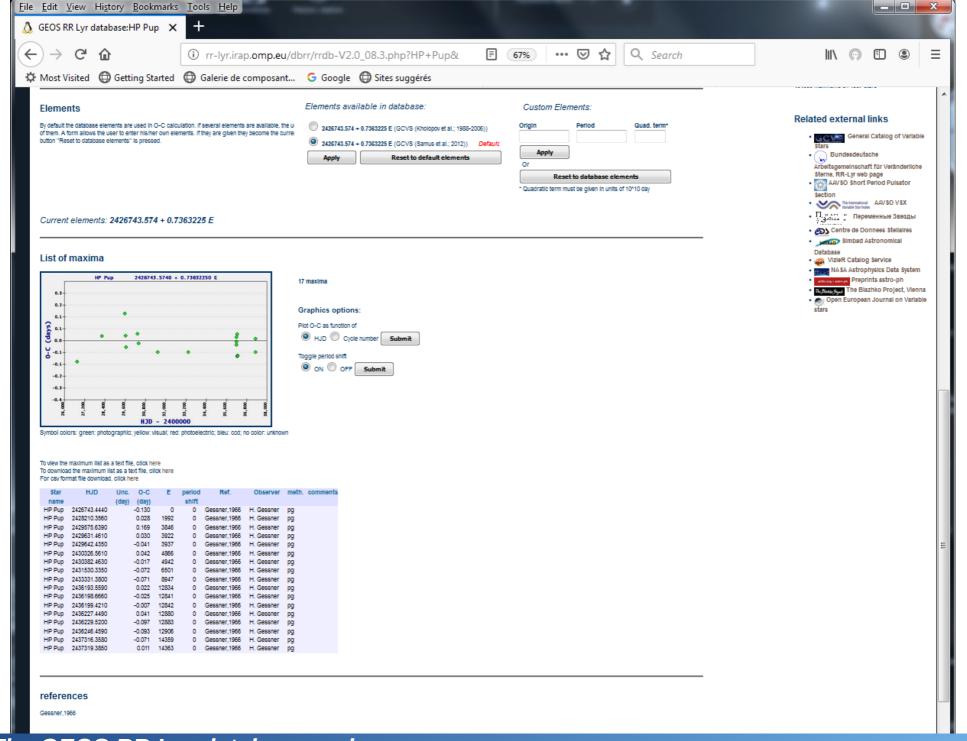


GCVS is updated irregularly at scale of several months,

Last modification: 09-Sep-2019

http://www.sai.msu.su/groups/cluster/gcvs/gcvs/





## Access to data and functions

#### Access to data

- Complete list of stars with maxima
- Access by constellation and GCVS names
- Access by coordinates
- Access by star name

HP Pup



- List observers
- Bibliography
- Statistics on RR Lyr maxima
- The Blazhko star catalog
- RR Lyr star spectroscopy

#### HP Pup:

- Compute ephemeris
- Search Period from maximum list
- GEOS RR Lyr wiki page
- \* SHITEAD
- AAVSO VSX record
- Dowload measurements from AAVSO
- AAVSO VVebObs Search Results

#### Lightcurve from AAVSO data

Enter JD of beginning and end of lightcurve and click on the " AA/SO lightcurve " button.

2458038.2340278

2458738.2340278

AAVSO lightcurve

#### Database numbers

#### Date of last updates

General RR Lyr catalog: 2019-06-09 17:02:01

(62313 entries)

Element table: 2019-08-19 18:21:52

(18755 entries)

Maximum table: 2019-09-06 18:53:53

104980 maximums on 4057 stars

#### Related external links

General Catalog of Variable

#### Stars

Bundesdeutsche

Arbeitsgemeinschaft für Veränderliche Sterne, RR-Lyr web page

AAVSO Short Period Pulsator

#### Section

AAVSO VSX

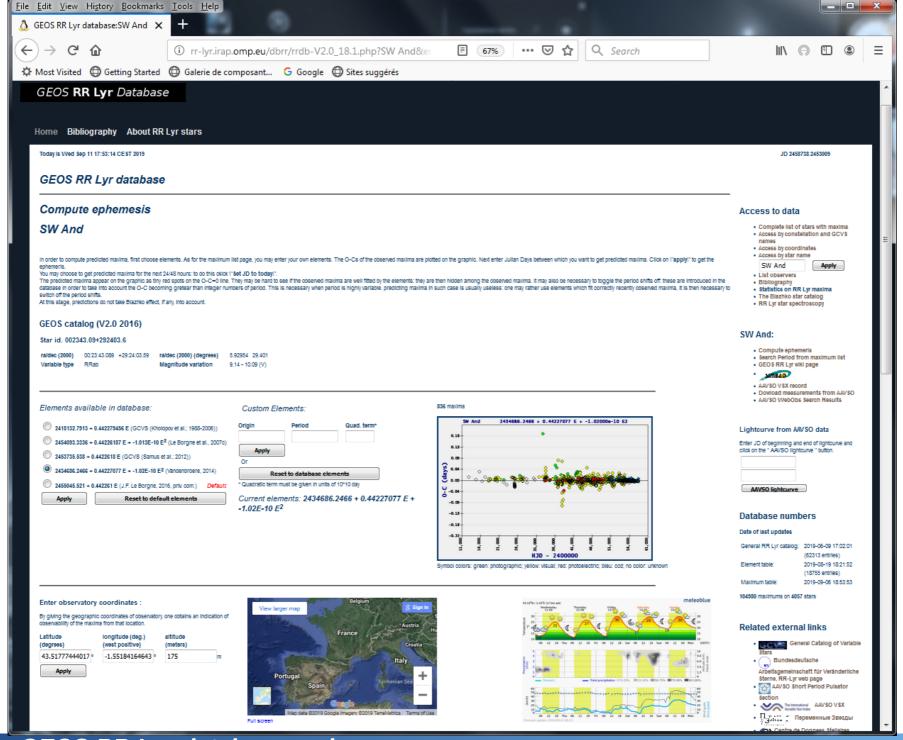
• Пописития Переменные Звезды

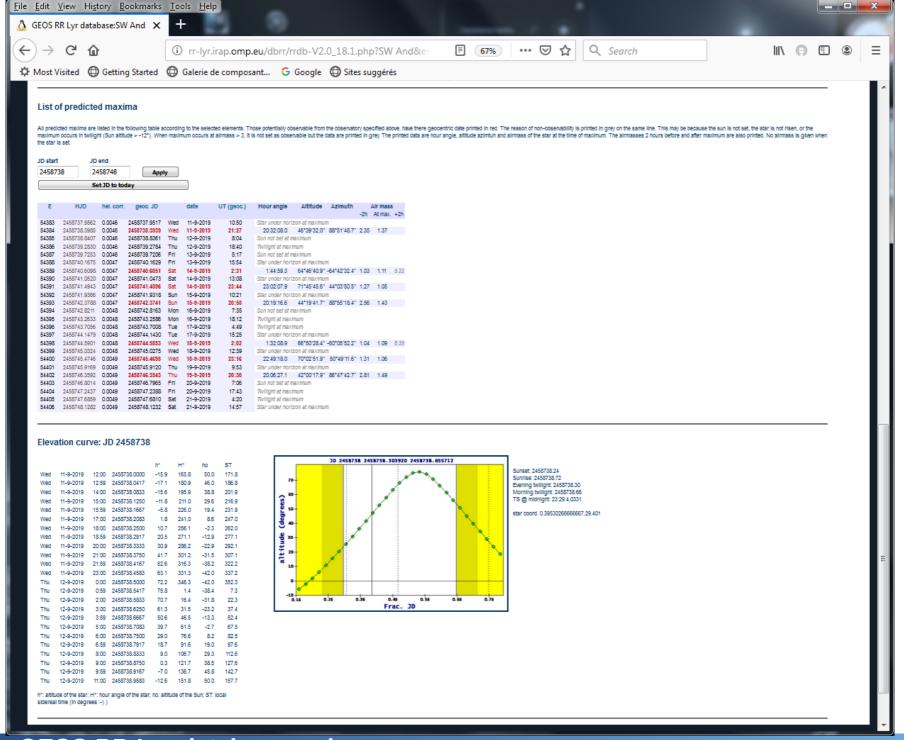
Centre de Donnees Stellaires

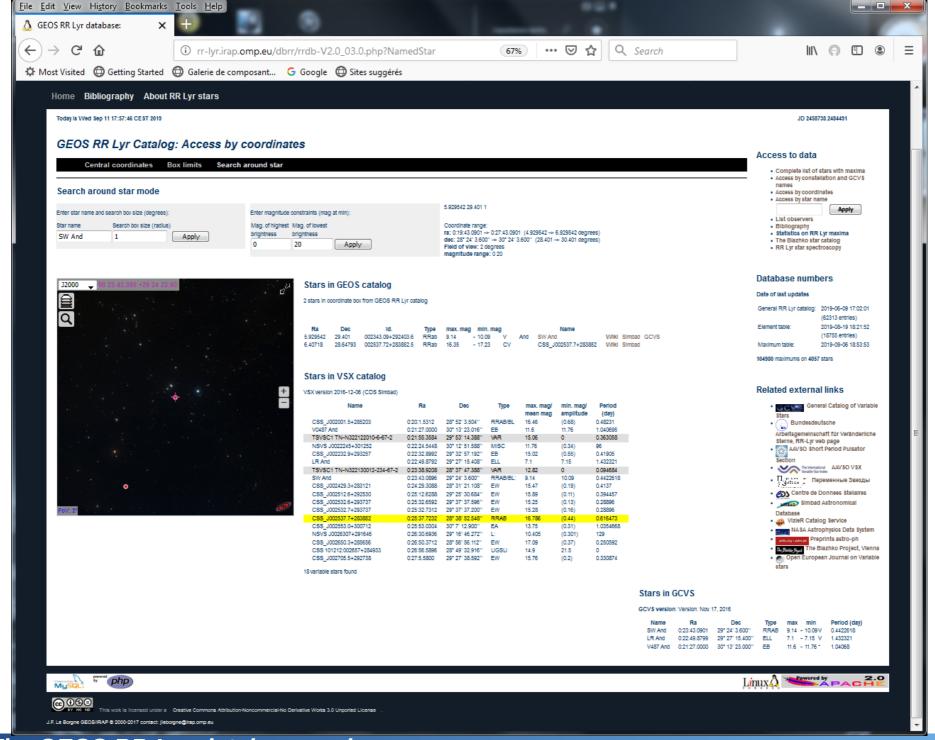
Simbad Astronomical

#### Database

- VizieR Catalog Service
- NASA Astrophysics Data System
- Preprints astro-ph
- The Blazhko Project, Vienna
- Open European Journal on Variable stars







## **GEOS RR Lyr Survey**

Started 2004

**Robots**: routine observations of maximums of bright RRab stars (magnitude at minimum <~13).

Aim: survey light curve variations at large time scales (>~10 years)

**Human observers**: Study of fainter understudied stars (magnitude at minimum range ~13-15).

Aim: refine or find period and find possible new Blazhko effects,

*Human observers*: Followup of RR Lyr itself with small dedicated instruments

## **Robotic Telescopes Tarot**

« Télescope à Action Rapide pour les Objets Transitoires »

Dedicated to followup of gamma ray bursts (M. Boër and A. Klotz) And observation of earth satellites for CNES.

Robotic and fast Mirror diameter 25 cm CCD cameras 2000x2000, field 2°x2°

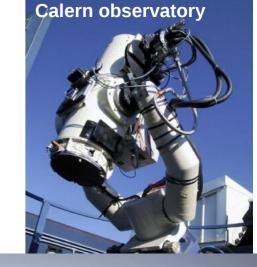
RR Lyr star survey is one of the additional programs

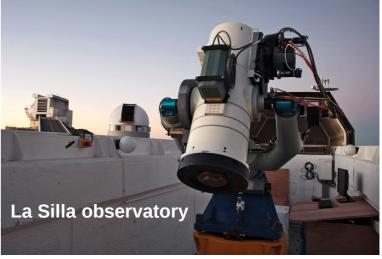
Calern, France: since 2004 9569 maximums 309 stars

La Silla, Chile: since 2006 8016 maximums 366 stars



Started 2019
Takahashi 15cm telescope
Celestron 28cm telescope



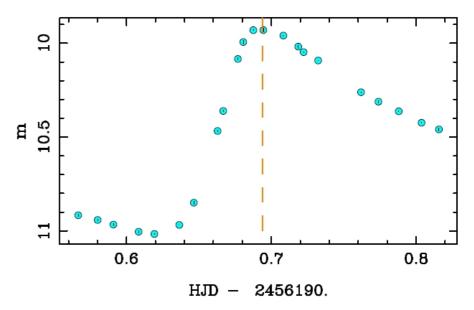


Automatic data reduction: bias, flat fields, photometry

## Routine followup of bright RR Lyr stars of type RRab

6 to 8 maxima scheduled on each telescope every night

Precision of ToMs: ~0.002 days (3mn)



#### Thursday 3 March 2016

			observation time (JD)		observation time (UT)		Maximum HJD			
JD	Star	N*	from	to	from	to	Predicted	Observed	O-C (day)	Observer/Telescope
2457451	BO Vir	16	0.409	0.508	21.8	0.2				Tarot - Calern
2457451		26	0.312	0.517	19.5	0.4				Tarot - Calern
2457451		20	0.315	0.434	19.6	22.4	2457451.343	2457451.348± 0.004	0.005	750
2457451	IM Leo	36	0.396	0.628	21.5	3.1	2457451.518	2457451.519± 0.002	0.001	Tarot - Calern
2457451	RV UMa	28	0.560	0.706	1.4	4.9	2457451.619	2457451.622± 0.002	0.002	Tarot - Calern
2457451	RX Leo	48	0.466	0.707	23.2	5.0	2457451.570	2457451.574± 0.003	0.004	Tarot - Calern
2457451	TW Boo	38	0.578	0.698	1.9	4.8	2457451.646	2457451.643± 0.001	-0.003	Tarot - Calern
2457451	V1361 Cen	2	0.899	0.900	9.6	9.6				Tarot - Chile
2457451	BI Cen	20	0.511	0.728	0.3	5.5	2457451.558	2457451.568± 0.003	0.011	Tarot - Chile
2457451	ET Hya	22	0.545	0.783	1.1	6.8	2457451.650	2457451.652± 0.003	0.003	Tarot - Chile
2457451	IU Car	24	0.595	0.799	2.3	7.2	2457451.687	2457451.694± 0.004	0.006	Tarot - Chile
2457451	TY Aps	22	0.644	0.860	3.5	8.6	2457451.670	2457451.672± 0.003	0.002	Tarot - Chile
2457451	V476 Vir	34	0.727	0.899	5.4	9.6	2457451.827	2457451.894± 0.005	0.067	Tarot - Chile
2457451	V689 Car	16	0.501	0.706	0.0	4.9	2457451.573	2457451.569± 0.004	-0.005	Tarot - Chile
2457451	X Crt	34	0.714	0.891	5.1	9.4				Tarot - Chile

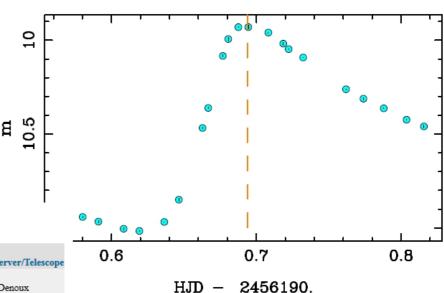
## Routine followup of bright RR Lyr stars of type RRab

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#### Saturday 17 August 2019

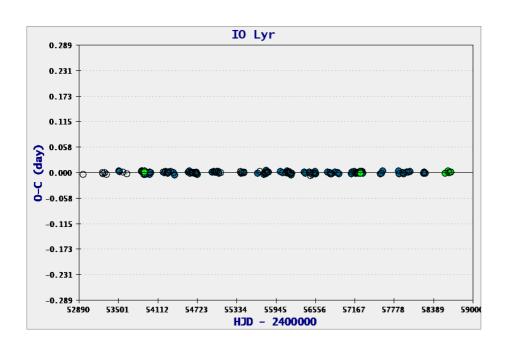
			1	(TP)	1	e arm		·		
-		374	observation time (JD)		observation time (UT)		Maximum HJD			
л	Star	N*	from	to	from	to	Predicted	Observed	O-C (day)	Observer/Telescope
2450742	DD D	22	0.270	0.667	21.1	4.0	2450712 542	2450712 527 0 005	0.016	F. D.
	BD Dra	33	0.378	0.667	21.1	4.0	2458713.543	2458713.527± 0.005		E. Denoux
	AO Peg	47	0.332	0.506	20.0	0.1	2458713.405	2458713.409± 0.003		A. Klotz
	CQ Lac	47	0.328	0.502	19.9	0.0	2458713.453	2458713.425± 0.002		A. Klotz
	DM And	41	0.513	0.637	0.3	3.3	2458713.567	2458713.570± 0.003		A. Klotz
	ES Peg	47	0.515	0.651	0.4	3.6	2458713.592	2458713.601± 0.003		A. Klotz
2458713	_	45	0.326	0.501	19.8	0.0	2458713.398	$2458713.392 \pm 0.002$		A. Klotz
2458713		42	0.511	0.635	0.3	3.2	2458713.598	2458713.610± 0.004		A. Klotz
2458713		36	0.566	0.667	1.6	4.0	2458713.623	2458713.621± 0.005		A. Klotz
2458713		30	0.397	0.488	21.5	23.7	2458713.479	2458713.478± 0.003	-0.001	A. Klotz
2458713		44	0.568	0.649	1.6	3.6				A. Klotz
2458713		36	0.509	0.633	0.2	3.2	2458713.591	2458713.592± 0.004	0.000	A. Klotz
2458713	V2630 Cyg	50	0.333	0.507	20.0	0.2				A. Klotz
	V509 Peg	47	0.330	0.504	19.9	0.1	2458713.426	2458713.428± 0.002	0.002	A. Klotz
2458713	RR Lyr	786	0.326	0.675	19.8	4.2	2458713.479	2458713.492± 0.003	0.013	A. Klotz
2458713	RT Dor	57	0.760	0.924	6.2	10.2	2458713.795	2458713.843± 0.003	0.049	F.J. Hambsch
2458713	BD Dra	33	0.378	0.668	21.1	4.0	2458713.543	2458713.525± 0.005	-0.018	E. Denoux
2458713	BD Dra	33	0.377	0.666	21.0	4.0	2458713.543	2458713.526± 0.005	-0.017	E. Denoux
2458713	PP UMa	23	0.373	0.663	21.0	3.9				E. Denoux
2458713	V740 Cep	33	0.380	0.669	21.1	4.1	2458713.547	2458713.537± 0.002	-0.010	E. Denoux
2458713	FW Peg	33	0.376	0.665	21.0	4.0	2458713.547	2458713.623± 0.006	0.075	E. Denoux
2458713	GV Peg	33	0.372	0.662	20.9	3.9				E. Denoux
2458713	UZ UMa	32	0.375	0.664	21.0	3.9				E. Denoux
2458713	NQ Dra	32	0.381	0.661	21.1	3.9	2458713.647	2458713.639± 0.005	-0.008	E. Denoux
	KQ UMa	13	0.374	0.663	21.0	3.9				E. Denoux
2458713	DD Aps	45	0.593	0.755	2.2	6.1	2458713.675	2458713.663± 0.005	-0.012	F.J. Hambsch
2458713	DI Aps	45	0.594	0.756	2.3	6.1				F.J. Hambsch
2458713	RW Oct	85	0.595	0.925	2.3	10.2	2458713.760	2458713.764± 0.003	0.004	F.J. Hambsch
2458713	VV Peg	57	0.395	0.545	21.5	1.1	2458713.484	2458713.505± 0.002	0.020	A. Klotz

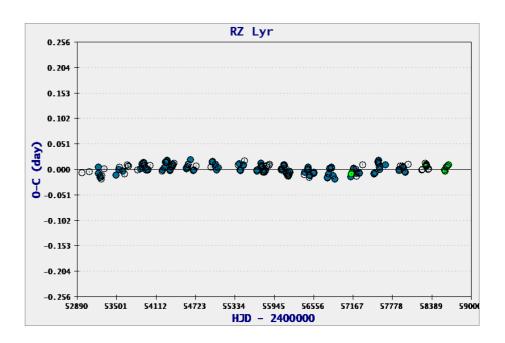


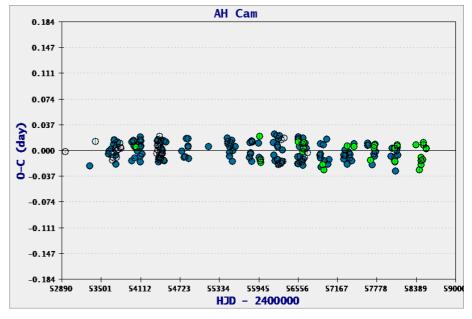
## Routine followup of bright RR Lyr stars of type RRab

6 to 8 maxima scheduled on each telescope every night

Precision of ToMs: ~0.002 days (3mn)







## **Human observers: Study of understudied stars**

Many RR Lyr satrs fainter than magnitude 13 at minimum are understudied.

Pulsation period and type are not accurately known

We do not know if there is a Blazhko effect

Observers use telescopes of 20 to 60 cm diameter

Observing plan: get complete folded light curve in a short time (about one week) and do it again several weeks or months later.

Variable number of amateur astronomers contribute to the program Presently from Europe and Japan Since 13 years, 25 astronomers have contributed, from Europe, Japan, Australia, USA and Namibia

~100 understudied RR Lyr stars observed

Several Blazhko effect discovered

### An example : CM Leo

Known as RRab proved to be an RRc

Observers: 2011-2015

Laurent Corp and collaborators, T60 Pic du Midi, France

Marco Nobile, 20cm Savosa, Switzerland Maurice Audejean, 30cm, Chinon, France

Also Francesco Fumagalli and Gisela Maintz (BAV)

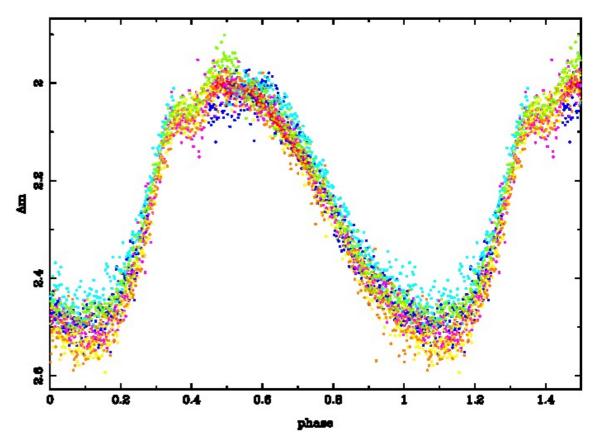
CM Leo 3357 mes.

Observed period: 0.366 jour

**GCVS: RRAB** 

Magnitude: 13.8-14.9 (B)

Period: 0.361732 jour



### **V568 Cas : A long period Blazhko effect**

GCVS: RR:

Magnitude 13.1-14.0 (p)

Period: 0.623 jour

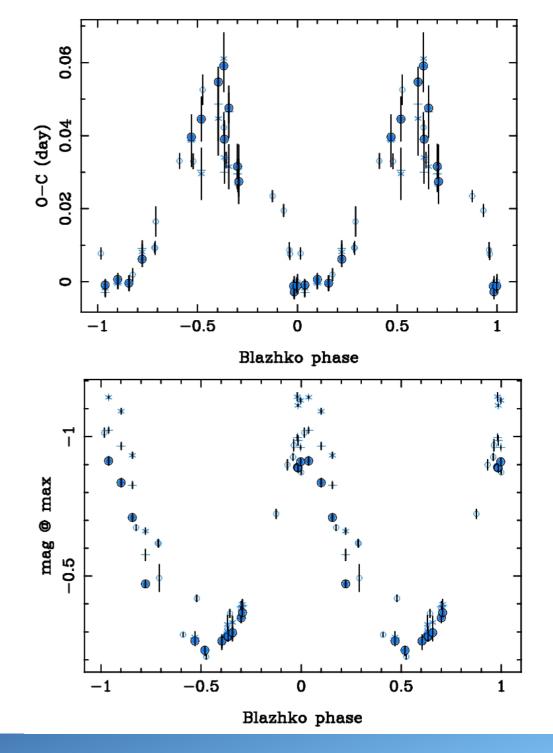
Observed period: 0.51404 days Blazhko effect period: 325 jours

Observers: 2008-2019

- Eric Denoux, Caussade, France, 28cm
- Mercè Correa, Freixinet Spain, 30cm
- Ramon Moliner, Mercè Correa, Florence

Libotte, Sabadell Spain, 50cm

- Kenji Hirosawa, Japan, 25cm



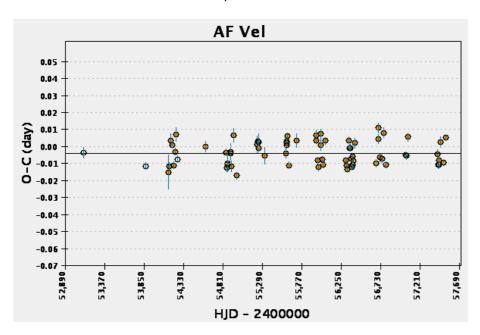
## Blazhko survey of southern RR Lyr stars

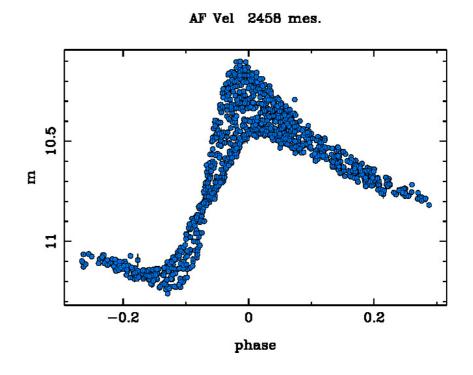
Observations in La Silla observatory over a decade has allowed to follow several hundreds of stars mostly understudied.

Also collaboration with Neil Butterworth in Australia in 2007-2011,

These observations have shown several new Blazhko effects among these stars.

**AF Vel**: Blazhko effect found by P. Wils and A. Sodor, 2005.





## Blazhko survey of southern RR Lyr stars

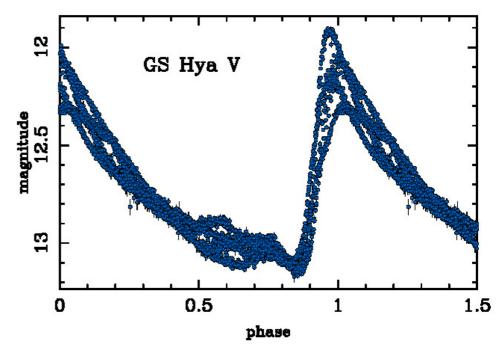
**New observations started in 2019** to confirm and characterize these new Blazhko effects from 2 remote observatories in Chile:

 Sadr observatory, Hacienda des étoiles, near La Serena observers: P. Traverse, J.L. Virlichie, S. Leclerc, P. Chatelain (Albireo association, France) telescopes: 355mm and 102mm

114 series (star-night), 17 stars

 Josch Hambsch's observatory, at Alain Maury's, San Pedro de Atacama telescopes: 400mm

307 series (star-night), 6 stars



Observations by Josch Hambsh

## Followup of the Blazhko effect of RR Lyr itself

### **RR Lyr**

magnitude 7-8, period 0.51 jour Blazhko effect period ~40 days.

#### Variable Blazhko effect:

Preston et al. (1965) observed disappearance of Blazhko effect in 1963.

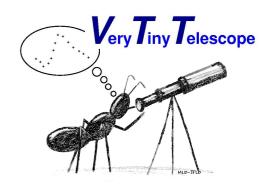
The phenomenon has a time scale of a few years,

Study of Variable Blazhko effect needs continuous observation during several years,

2008-2019:

2 VTTs: 1303 nights of observations 507 maxima

In total with other observers (+literature) 734 maxima (559 different)





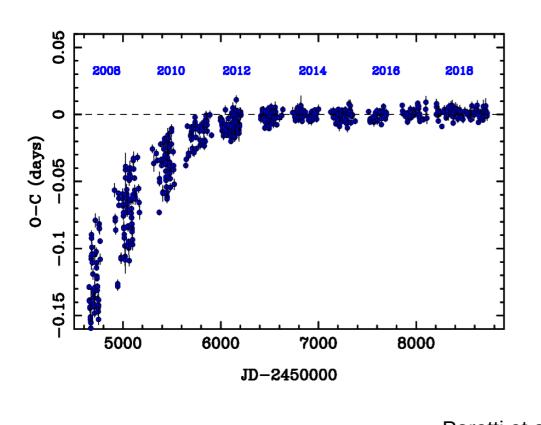
Instrumentation : VTT, Design : Alain Klotz

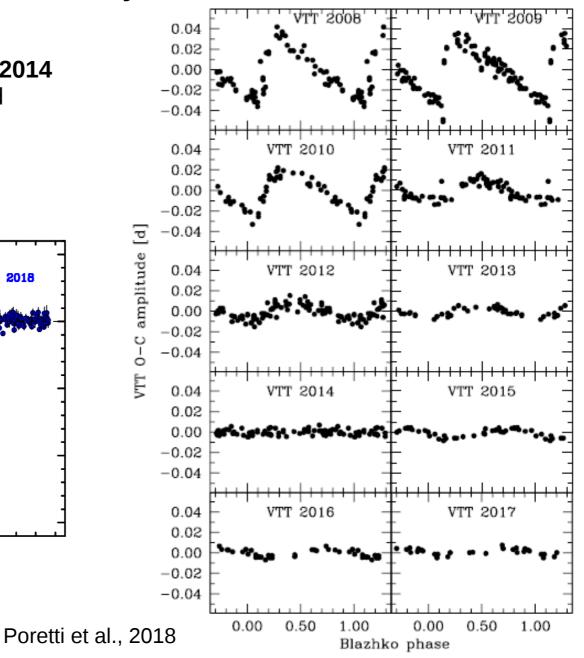
Camera CCD Audine kaf400 Photo lens 135mm F/2.8 Mount HEQ5 GOTO controlled by computer

Winter 2018-2019: no interruption, observations in Norway by Jan Qvam and Ø. Midtskogen

## Followup of the Blazhko effect of RR Lyr itself

Disappearance of Blazhko effect in 2014 as in 1963 (Le Borgne et al. 2014) and Restart in 2015
But still very faint in 2019





## Conclusions

To find new things on RR Lyr stars, no need of giant telescopes.

**But** need to observe a lot (short time sampling, minutes), during a long time (years): this is slow science!

Solution: collaborative work and use of automated/robotic telescopes.

Thank you